

STiCM

Select / Special Topics in Classical Mechanics

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STiCM Lecture 40:

The Scope, and Limitations, of Classical Mechanics

The Scope, and Limitations, of Classical Mechanics

widely accepted, used for a long time;
traditional in style or idea



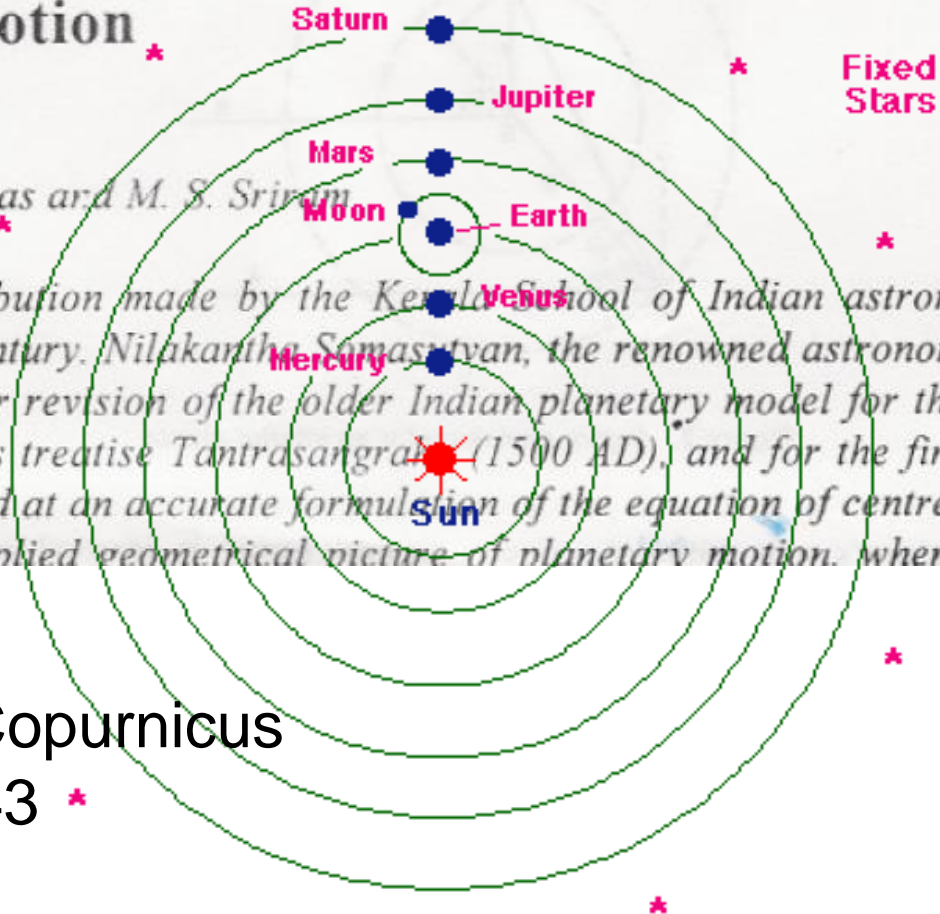
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Modification of the earlier Indian planetary theory by the Kerala astronomers (c. 1500 AD) and the implied heliocentric picture of planetary motion

K. Ramasubramanian, M. D. Srinivas and M. S. Srinivasan

We report on a significant contribution made by the Kerala School of Indian astronomers to planetary theory in the fifteenth century. Nilakantha Somashtyan, the renowned astronomer of the Kerala School, carried out a major revision of the older Indian planetary model for the interior planets, Mercury and Venus, in his treatise *Tantrasangraha* (1500 AD), and for the first time in the history of astronomy, he arrived at an accurate formulation of the equation of centre for these planets. He also described the implied geometrical picture of planetary motion, where the five



Nicolus Copernicus
1473-1543 *

<http://www.physics.iitm.ac.in/~labs/amp/kerala-astronomy.pdf>

Albert Einstein: “We owe a lot to Indians, who taught us how to count, without which no worthwhile scientific discovery could have been made.”

ARYABHATTA (in 5th century) introduced new concepts: sphericity of the earth, rotation about its axis, revolution around the sun, explanation of eclipses..... estimated length of the year.....

BRAHMAGUPTA (in 7th century) estimated the circumference of the earth to be around 5000 yoganans which in today's units is close to the correct value as we know it now....

Central problem in ‘Mechanics’: How is the ‘mechanical state’ of a system described, and how does this ‘state’ evolve with time?

‘position’ and ‘velocity’: both needed
to specify the mechanical state of a system?

The mechanical state of a system is characterized by its position and velocity, (q, \dot{q})

or, position and momentum, (q, p)

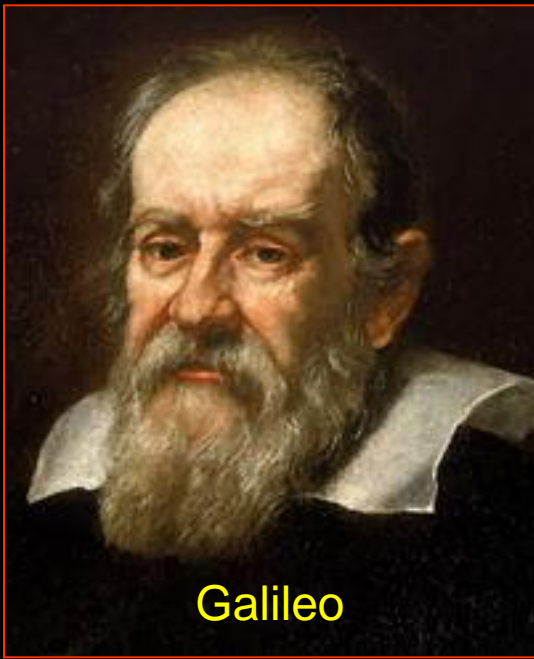
Or, equivalently by their well-defined functions:

$L(q, \dot{q})$: Lagrangian

$H(q, p)$: Hamiltonian

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Galileo



Newton

$$(q, \dot{q})$$

$$\vec{F} = m\vec{a}$$

Linear Response.

Principle of causality.

Principle of
Variation

$$L(q, \dot{q})$$

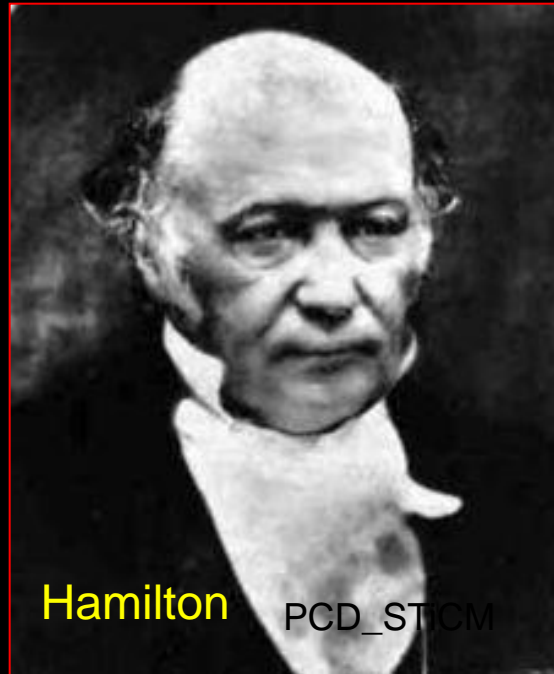
$$H(q, p)$$

$$\int_{t_1}^{t_2} L(q, \dot{q}, t) dt$$

$$\frac{\partial L}{\partial q} - \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}} \right) = 0$$



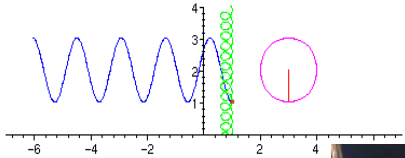
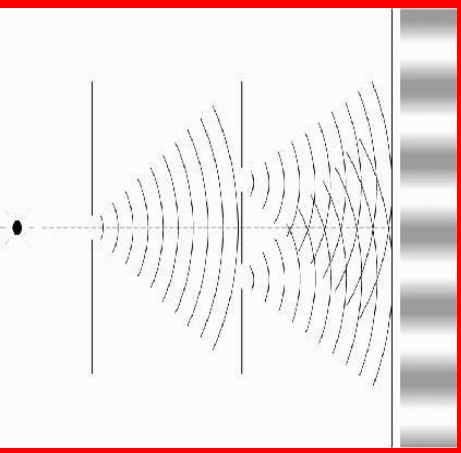
Lagrange



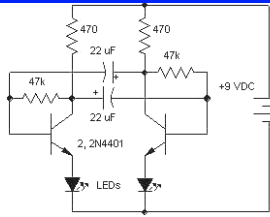
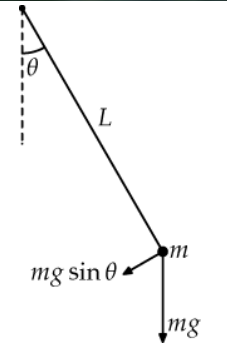
Hamilton

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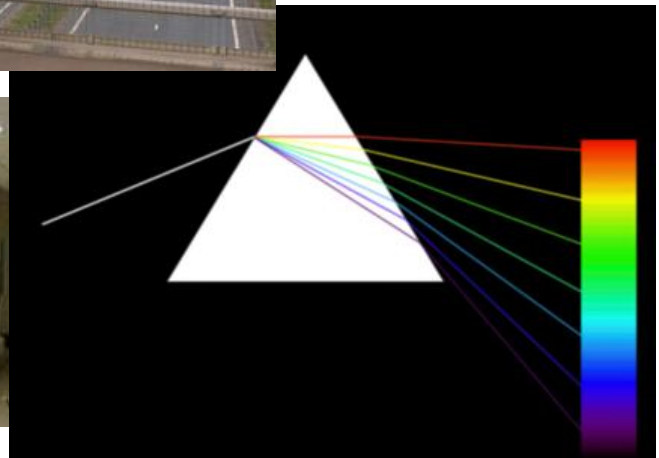
$$\dot{q} = \frac{\partial H}{\partial p}, \dot{p} = -\frac{\partial H}{\partial q_k}$$



Oscillations.
 Small oscillations.
 SHM.
 Driven and damped oscillator.
 Resonance,
 Quality factor.
 Waves.



Flip-Flop Square Wave oscillator



$$\frac{c}{v_{\varphi}} = n = \frac{\lambda_{\text{vac}}}{\lambda_{\varphi}}$$

$$n_r = n_r(\omega)$$

*My heart leaps up when I behold
A rainbow in the sky:
So was it when my life began;
So is it now I am a man;
So be it when I shall grow old,
Or let me die!...
- William Wordsworth*

R.I. of water
for red is
~1.331

R.I. of water
for blue is
~1.343

Questions:

1. Why is the **red outside** and **blue inside**?
2. Which part of this picture is the brightest, and why?



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Rainbow, seen from the 'Maid of the Mist' ride at the
Niagara Falls, U.S.A., 18th July, 2009. - pcd

Are the conservation principles consequences of the laws of nature? Or, are the laws of nature the consequences of the symmetry principles that govern them?

Until Einstein's special theory of relativity, it was believed that conservation principles are the result of the laws of nature.

Since Einstein's work, however, physicists began to analyze the conservation principles as consequences of certain underlying symmetry considerations, enabling the laws of nature to be revealed from this analysis.

Instead of introducing Newton's III law as a *fundamental principle*,
we deduced it (in Unit 1) from symmetry / invariance.

This approach places SYMMETRY *ahead of* LAWS OF NATURE.

It is this approach that is of greatest value to contemporary physics. This approach has its origins in the works of Albert Einstein, Emmily Noether and Eugene Wigner.



(1879 – 1955)



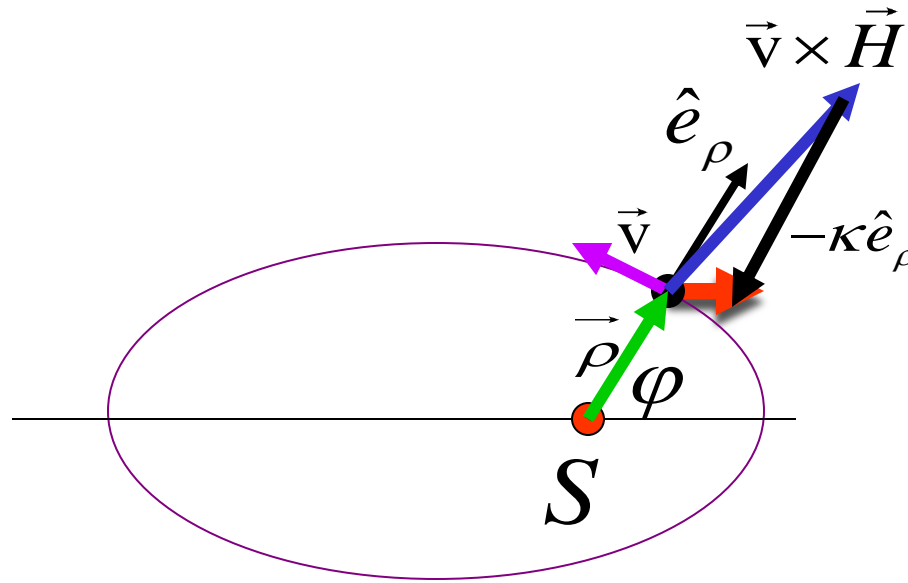
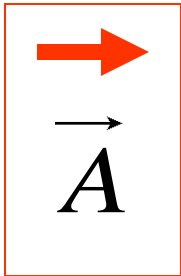
(1882 – 1935)



(1902 – 1995)

Laplace Runge Lenz Vector

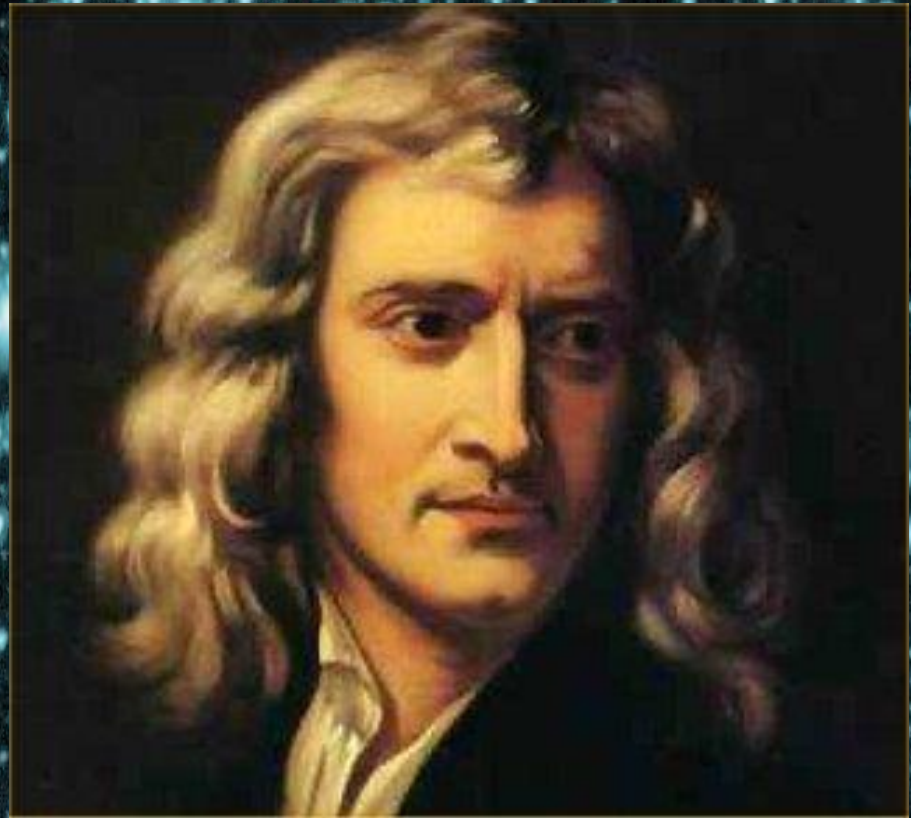
is constant for a strict $\frac{1}{r}$ potential.



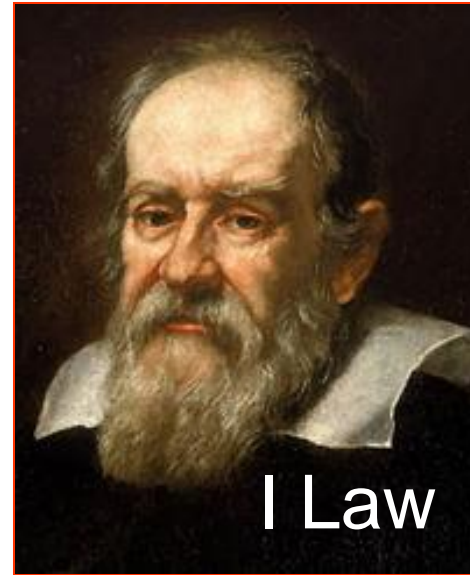
The $\odot \vec{H}$
(specific)
angular
momentum
vector is
out of the
plane of
this figure,
toward us.

$$\vec{A} = \left(\vec{v} \times \vec{H} \right) - \kappa \hat{e}_\rho$$

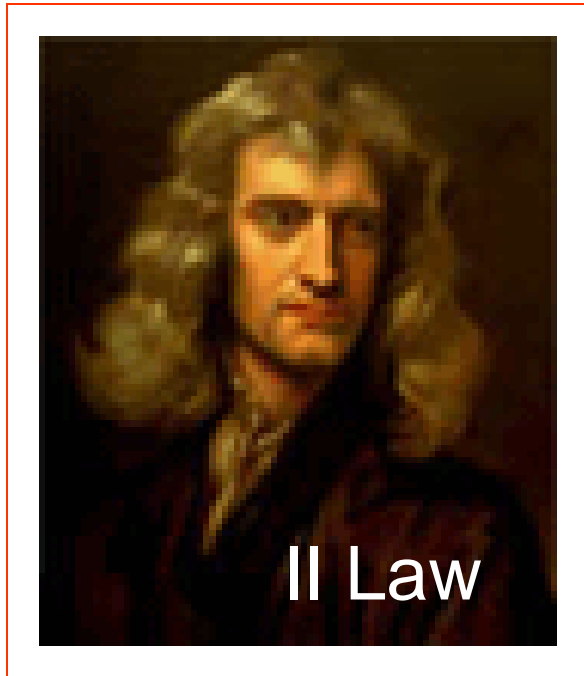
**Just where is the inertial frame?
Newton envisaged the inertial
frame to be
located in
deep space,
amidst
*distant stars.***



Galileo's experiments that led him to the law of inertia.



Galileo Galilei
1564 - 1642

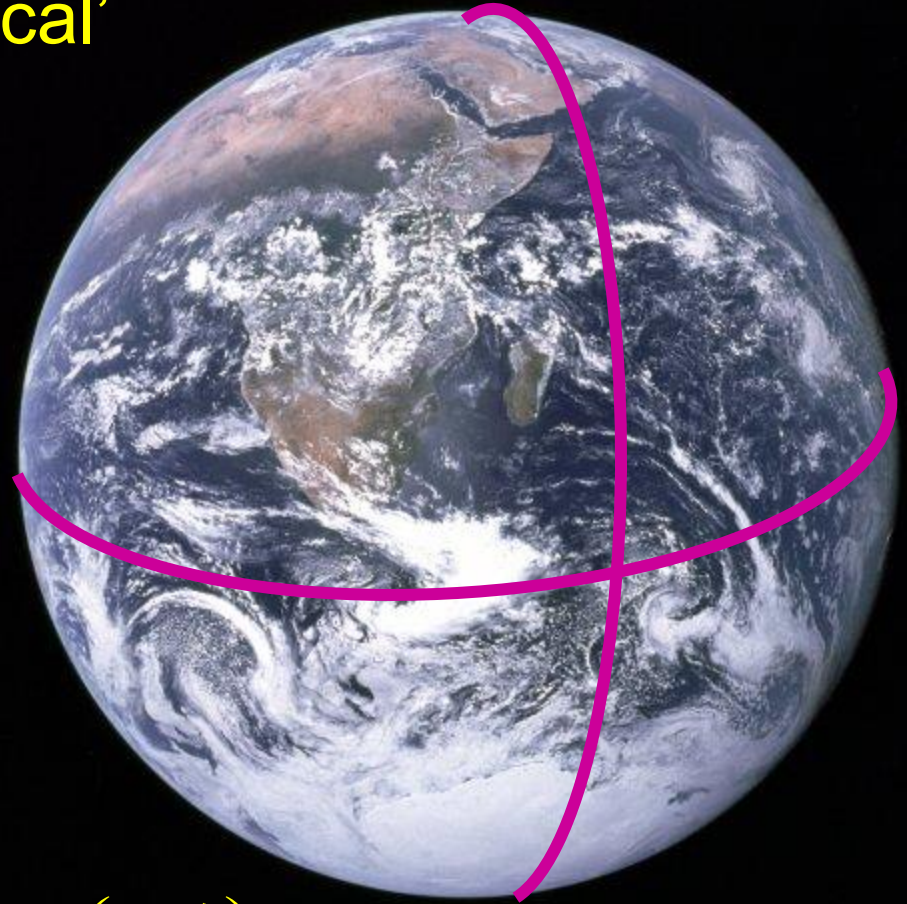


$$\vec{F} = m\vec{a} \quad \text{Linear Response.}$$

Effect is proportional to the *Cause*

Principle of causality.

3 'definitions' of the 'vertical'



$$m \left(\frac{d^2}{dt^2} \right)_R \vec{r} = m \left(\frac{d^2}{dt^2} \right)_I \vec{r} - m \left(\frac{d\vec{\omega}}{dt} \right)_R \times \vec{r} - 2m\vec{\omega} \times \left(\frac{d}{dt} \right)_R \vec{r} - m\vec{\omega} \times (\vec{\omega} \times \vec{r})$$



$$\begin{pmatrix} -x \\ y \\ z \end{pmatrix} \leftarrow \begin{pmatrix} x \\ y \\ z \end{pmatrix}$$

REFLECTION

$$\begin{pmatrix} -1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} x \\ y \\ z \end{pmatrix} = \begin{pmatrix} -x \\ y \\ z \end{pmatrix}$$

Left \longleftrightarrow Right

Top $\overset{?}{\longleftrightarrow}$ Bottom

$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & -1 \end{pmatrix} \begin{pmatrix} x \\ y \\ z \end{pmatrix} = \begin{pmatrix} x \\ y \\ -z \end{pmatrix}$$

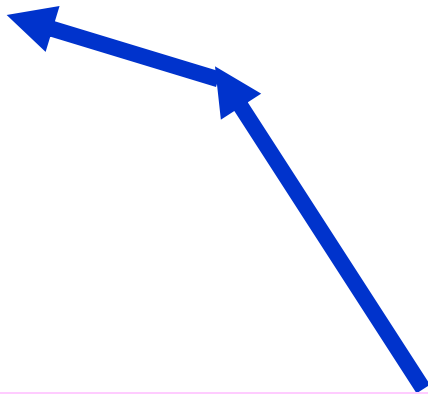
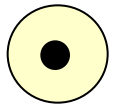
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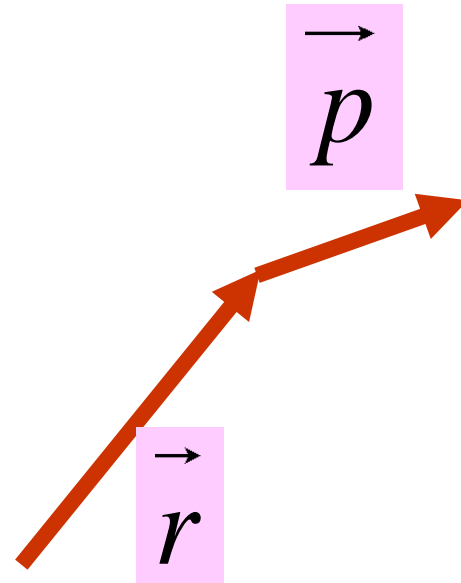
$$\vec{C} = \vec{A} \times \vec{B}$$

$$\vec{l} = \vec{r} \times \vec{p}$$

angular momentum



mirror



\vec{l} right-hand-cross-product

$$= \vec{r}_{image} \times \vec{p}_{image}$$



$$\vec{l} = \vec{r} \times \vec{p}$$

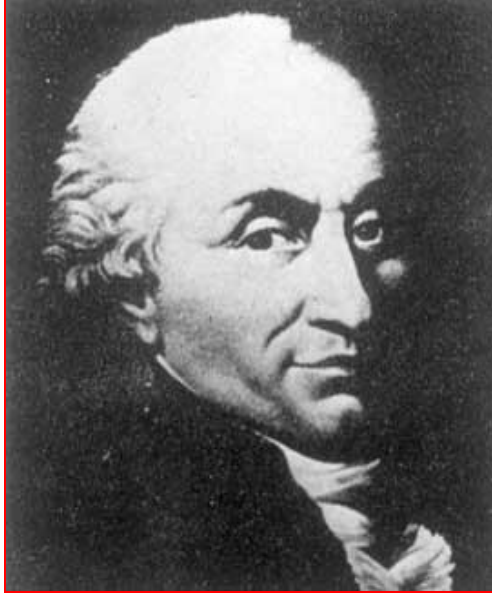
Galilean Relativity



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$$\frac{d\vec{r}}{dt} - \vec{u}_c = \frac{d\vec{r}'}{dt}$$

Speed of light ?



Charles Coulomb
1736-1806



Andre Marie
Ampere
1775-1836

Michael Faraday
1791-1867

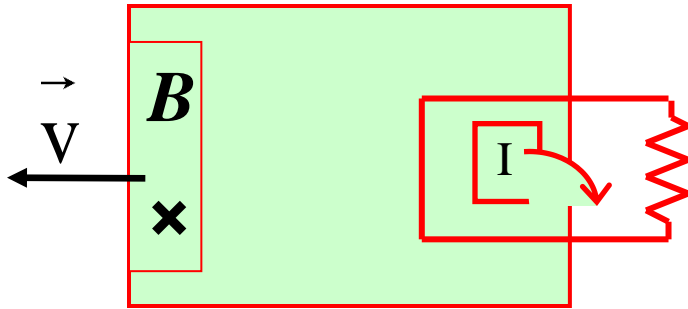


Carl Freidrich
Gauss
1777-1855



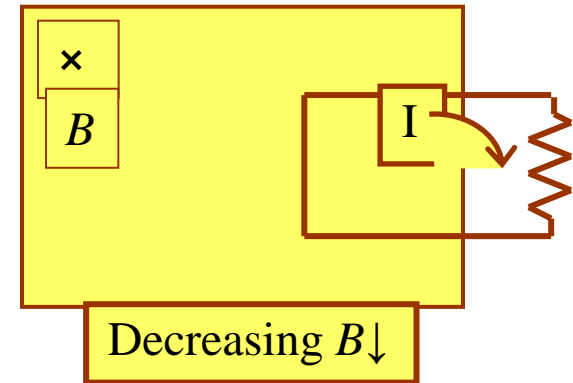
..... other
developments in
Physics

Faraday's experiments



Current: identical!

Strength of B decreased.
Nothing is moving,
but still, current seen!!!



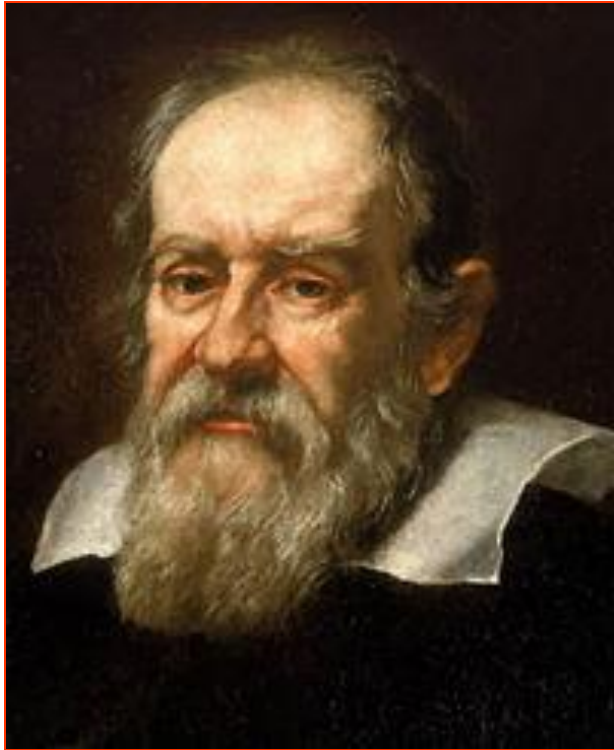
$$I \propto \frac{dB}{dt}$$

Einstein:

Special Theory of Relativity

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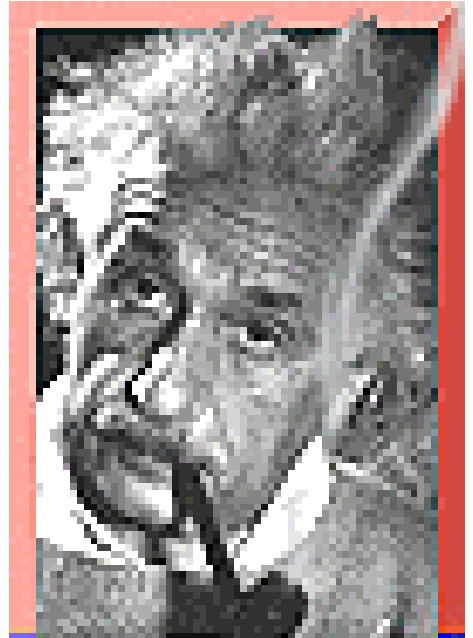
Galilean & Lorentz Transformations. Special Theory of Relativity.



Galileo Galilei
1564 - 1642



Hendrik Antoon Lorentz
1853-1928



**Smoking is
injurious to
health!**

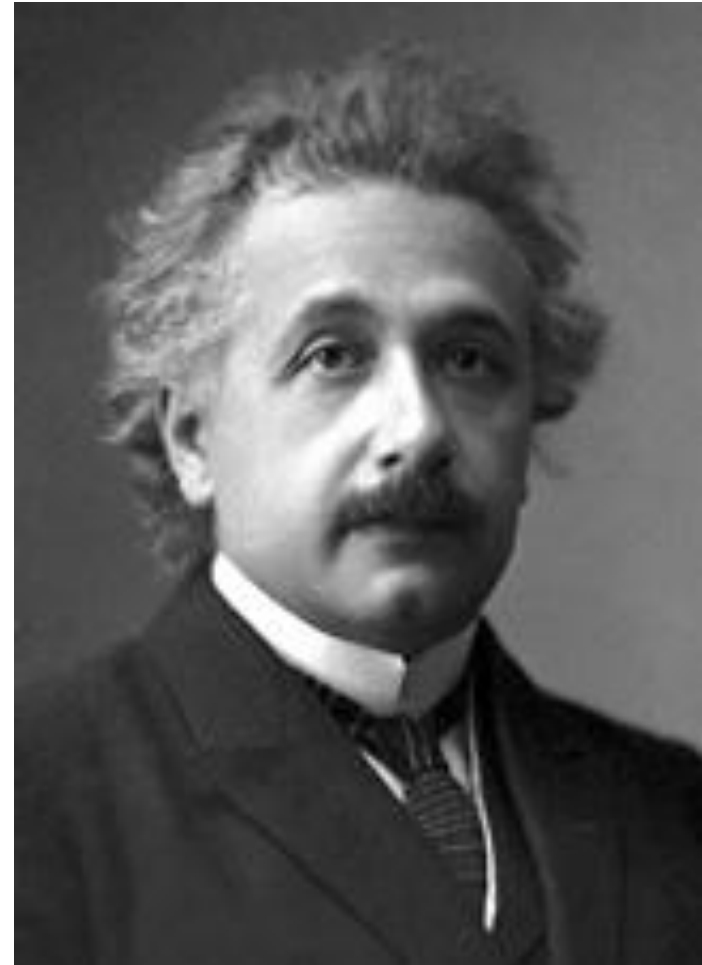
Albert Einstein
1879-1955

Classical Electrodynamics and the Special Theory of Relativity.

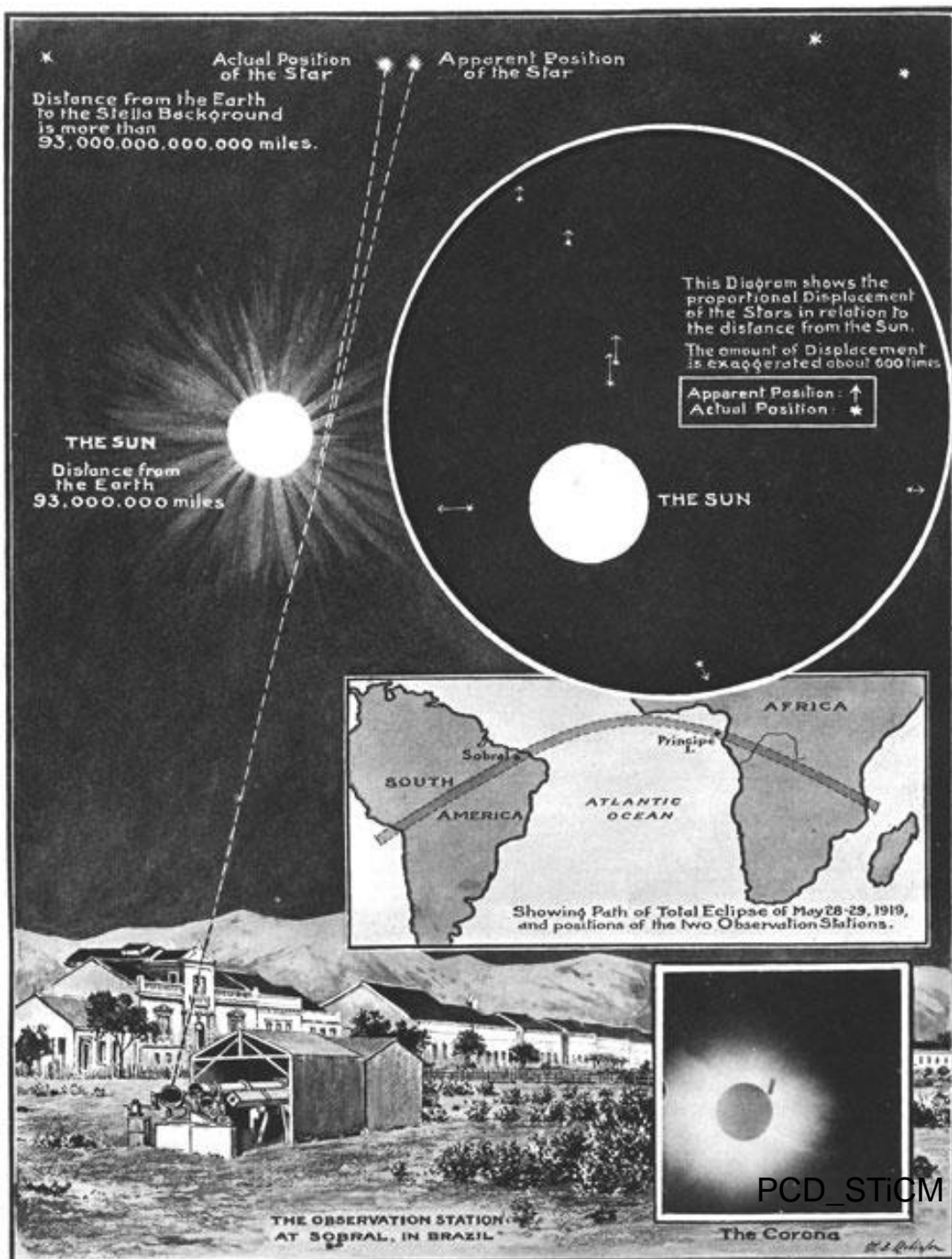


Maxwell

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Einstein



Arthur Eddington

http://denisdutton.com/einstein_eddington.htm

Observation station at SOBRAL, BRAZIL



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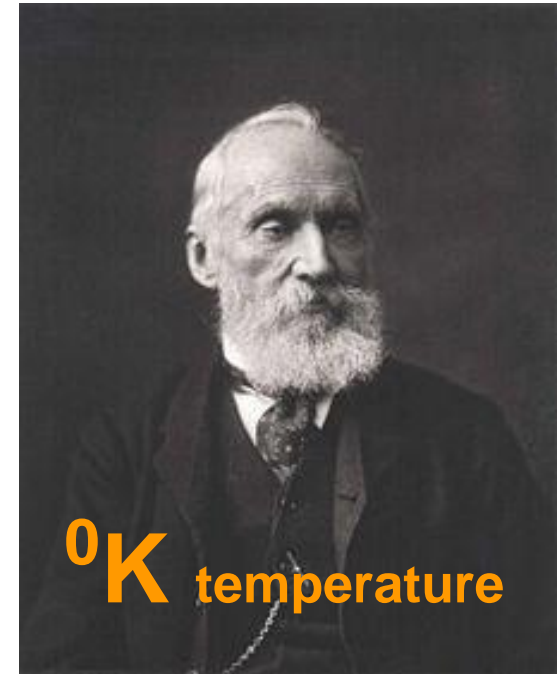
Johann
Carl
Friedrich
Gauss
1777 - 1855



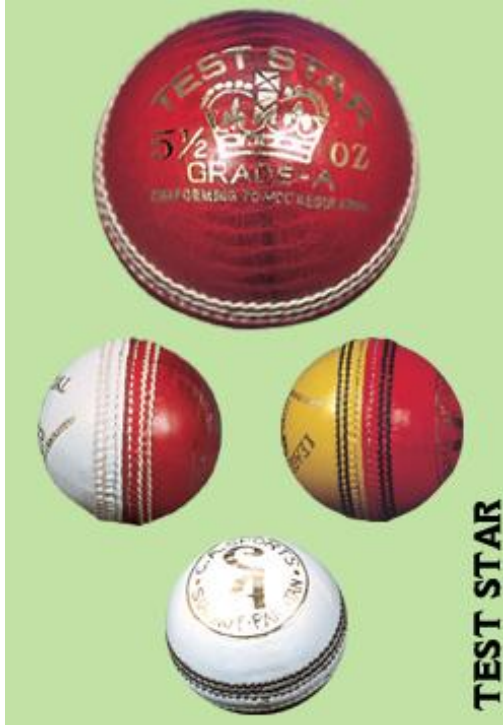
George G. Stokes
1851

George Gabriel
Stokes
(1819–1903)

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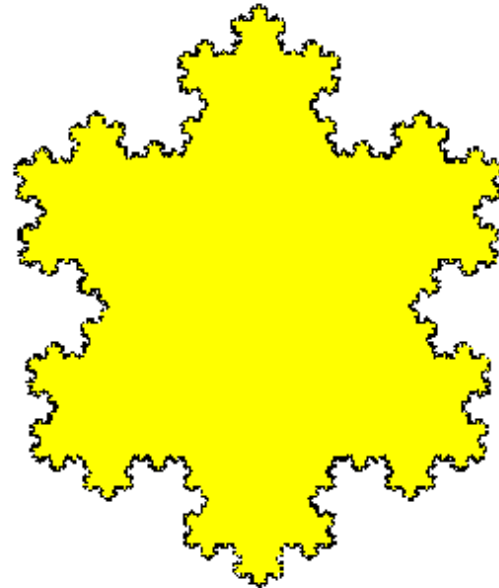
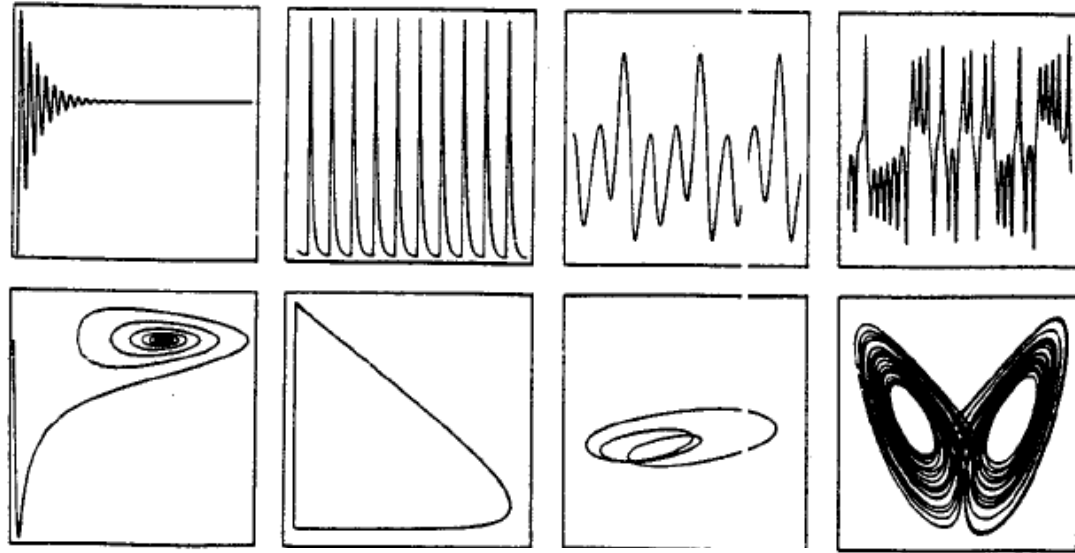
William Thomson,
1st Baron Kelvin
(1824-1907)



Ishant Sharma
Swing Bowling

Difference between the rough and shiny surface of a white ball is much more, and hence swings more!

'Chaos', bifurcation, strange attractors, fractals, self-similarity, Mandelbrot sets.



Principal premise of “classical mechanics”

mechanical system is characterized by position and velocity/momentum,

simultaneously and accurately.

Central problem in 'Mechanics':

How is the 'mechanical state' of a system described, and how does this 'state' evolve with time?

- Formulations due to Galileo/Newton,
- Lagrange and Hamilton.

(q,p) : How do we get these?

MEASUREMENT !

New approach required !

'New approach' is not required on account of the Heisenberg principle!

Rather,

the measurements of q and p are not compatible....

It is rather the HEISENBERG PRINCIPLE that

RESULTS as an expression of this incompatibility!

..... so how could one describe the

mechanical state of a system by (q,p) ?

Mechanical State:
State vectors in Hilbert Space

Characterize? Labels?

“Good” quantum
numbers/labels

Measurement: C.S.C.O. >

Complete Set of Commuting Operators

Complete Set of Compatible Observables

$$i\hbar \frac{\partial}{\partial t} | \rangle = H | \rangle$$

Schrödinger Equation

Evolution of the
Mechanical
State of the system

Quantization! state vector: $|\ \ \rangle$

dynamical variables: operators

$$A |\ \ \rangle \rightarrow |\ \text{label ?} \ \rangle$$

$$|\ \text{new vector} \ \rangle = \alpha |\ \text{old vector} \ \rangle$$

eigenvalue equation

$$A |\ \ \rangle = a |\ \text{label ?} \ \rangle$$

$$A |\ a \ \rangle = a |\ a \ \rangle$$

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$$A |\ a \ \rangle = a |\ a \ \rangle$$

eigenvalue equation

$|\text{label}(s)?\rangle$

 label

$$A | \rangle = a | a \rangle$$

$$A | a \rangle = a | a \rangle$$

eigenvalue equation

Measurement: system 'collapses' into its eigenstate

What *else* can be measured ?

C.S.C.O.

Complete Set of Compatible Observables

Complete Set of Commuting Operators

$$B | a, b \rangle = b | a, b \rangle \quad [A, B]_- = AB - BA$$

CSCO: $\{A, B, C, \dots\}$



Erwin Schrödinger

1887 - 1961



Werner Heisenberg

1901 - 1976

..... Any questions ?

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